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S/048/60/024/009/007/015  
B013/B063

24.6720

AUTHORS: Malysheva, T. V., Khotin, B. A., Lavrukhina, A. K.,  
Kryukova, L. N., Murav'yeva, V. V.

TITLE: Investigation of New Neutron-deficient Platinum Isotopes /9

PERIODICAL: Izvestiya Akademii nauk SSSR. Seriya fizicheskaya, 1960,  
Vol. 24, No. 9, pp. 1109-1112

TEXT: The authors studied neutron-deficient platinum isotopes which were formed by disintegration of gold induced by 660-Mev protons. The platinum isotopes were studied by spectrum analysis of conversion electrons and on the basis of "genetic" relations. The platinum and iridium fractions of high specific activity were separated from 1-2 g of gold bombarded on the synchrocyclotron of the OIYaI (Joint Institute of Nuclear Research). A series of experiments was carried out, and a method was proposed for the carrier-free separation of platinum and iridium isotopes. This method is based on the separation of their chloride complex compounds by means of anion exchange (Fig. 1). The results of the spectrum analysis of conversion electrons of the iridium fraction are given in Ref. 6. The spectrum of

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Investigation of New Neutron-deficient  
Platinum Isotopes

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the Pt conversion electrons was measured by means of a magnetic, spiral  $\beta$ -spectrometer having a resolution of 0.5%. The experimental spectrum is shown in Fig. 2. The results of measurement of the conversion lines are given in a table. The half-life of the Pt isotopes was determined by separating the daughter iridium from the Pt fraction at regular intervals during one to four hours (Fig. 3). The total half-life of  $Pt^{186}$  and  $Pt^{187}$  was calculated from the activity of  $Ir^{186}$  ( $T = 15$  hours) and  $Ir^{187}$  ( $T = 13$  hours) to be  $2.5 \pm 0.5$  hours. This is in agreement with the data of Ref. 6. The half-life of Pt was calculated from the activity of the daughter iridium having a half-life of three hours to be  $2.6 \pm 0.6$  hours. In accordance with Ref. 6, this is the half-life of the new isotope  $Pt^{184}$ . There are 3 figures, 1 table, and 10 references: 2 Soviet, 2 Canadian, and 1 German.

ASSOCIATION: Institut geokhimii i analiticheskoy khimii im. V. I. Vernad-  
skogo Akademii nauk SSSR (Institute of Geochemistry and  
Analytical Chemistry imeni V. I. Vernadskiy of the Academy of  
Sciences USSR). Nauchno-issledovatel'skiy institut yadernoy  
fiziki Moskovskogo gos. universiteta im. M. V. Lomonosova  
(Scientific Research Institute of Nuclear Physics of Moscow  
State University imeni M. V. Lomonosov)

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S/048/60/024/009/008/015  
B063/B063

26.1640

AUTHORS:

Lavrukhina, A. K., Kolesov, G. M., Tan Syao-yen

TITLE:

Neutron-deficient Isotopes of Rare-earth Elements of the Cerium Group (Products of the Disintegration of Europium With 660-Mev Protons)

PERIODICAL:

Izvestiya Akademii nauk SSSR. Seriya fizicheskaya, 1960, Vol. 24, No. 9, pp. 1113-1118

TEXT: The present paper gives the results of an investigation of the production cross sections of rare-earth isotopes which are produced by the disintegration of europium with 660-Mev protons. A target of 26.5 mg of europium oxide (99.8%  $\text{Eu}_2\text{O}_3$ ) was irradiated with the inner 660-Mev proton beam of the synchrocyclotron of OIYAI (Joint Institute of Nuclear Research) for 50 min. The fractions were identified from the position of the peaks in the chromatogram (Fig. 1) and the active rare-earth isotopes from the half-life and the radiant energy (Table 1). The yields of isotopes produced by the disintegration of europium are shown in Fig. 2

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Neutron-deficient Isotopes of Rare-earth  
Elements of the Cerium Group (Products of the  
Disintegration of Europium With 660-Mev Protons)

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Table 2 gives exact data on the composition of the isotopes produced by total decay of all radioisotopes. Finally, the authors briefly describe the separation of the predicted new isotope  $\text{Pr}^{134}$  (Ref. 7). This isotope originates from  $\text{Pr}^{134}$  whose half-life was estimated from the activity of  $\text{Ce}^{134}$  separated from the daughter fractions to be  $\sim 1$  hour. The gamma spectrum of the daughter cerium was taken on a 100-channel scintillation gamma spectrometer by Yu. A. Surkov and V. G. Karpushin (Fig. 5). V. V. Malyshev participated in the experiments. There are 5 figures, 2 tables, and 10 references: 8 Soviet.

ASSOCIATION: Institut geokhimii i analiticheskoy khimii im. V. I. Vernadskogo Akademii nauk SSSR (Institute of Geochemistry and Analytical Chemistry imeni V. I. Vernadskiy of the Academy of Sciences USSR)

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S/048/60/024/009/009/015  
B013/B063

24.6720  
AUTHORS:

Surkov, Yu. A., Chernov, G. M., Lavrukhina, A. K.,  
Khromchenko, Z. V.

TITLE:

Investigation of Neutron-deficient Osmium Isotopes

PERIODICAL:

Izvestiya Akademii nauk SSSR. Seriya fizicheskaya, 1960,  
Vol. 24, No. 9, pp. 1119-1123

TEXT: The present paper gives the results of an investigation of neutron-deficient osmium isotopes on the synchrocyclotron of OIYaI (Joint Institute of Nuclear Research). The osmium isotopes were produced by bombarding ~0.2 g of gold with 660-Mev protons for 1-2 hours. The purity of the separated elements was radiochemically checked. A 100-channel scintillation gamma spectrometer and  $\beta$ - and  $\gamma$ -counters were used to analyze the beta and gamma emission of the nuclear reaction products. Fig. 1 shows the gamma spectrum of the osmium isotopes obtained. The authors identified  $Os^{182}$ ,  $Os^{183}$ ,  $Os^{183*}$ , and  $Os^{185}$  which had a half-life of at least 10 hours. In addition, the spectrum showed an intense line,  $E_{\gamma} = 230$  kev, which had

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Investigation of Neutron-deficient Osmium Isotopes

a half-life of  $\sim 2.7$  hours. Control experiments indicated the existence of a new osmium isotope having a half-life of three hours. This was confirmed by the study of the daughter osmium (Figs. 3 and 4) and the daughter rhenium (Fig. 5). This neutron-deficient isotope is assumed to be  $\text{Os}^{181}$  and has a half-life 2.7 hours. By capture of the orbital electron it is converted into  $\text{Re}^{181}$ . 230-kev gamma quanta are emitted during this conversion. Fig. 2 shows the descending curve of the activity sum of Os, which was measured by an end-window counter. It confirms the correctness of the identification of the isotopes. As there are now only few data available on neutron-deficient Ir, Os, and Re isotopes, the conclusions drawn from the results obtained require further confirmation. The agreement of these results with experiments recently carried out with protons of 10-80 Mev (Ref. 1) indicates that the above-mentioned identification is correct. The results further indicate the existence of the isotope  $\text{Ir}^{183/19}$  which is formed by the decay of the two isomers  $\text{Os}^{183}$  and  $\text{Os}^{183\text{m}}$ . Furthermore, a 23-min activity of osmium ( $E_\gamma = 170$  kev) was observed during the experiments. However, the 23-min isotope may well be an isomer

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Investigation of Neutron-deficient Osmium  
Isotopes

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of some unidentified osmium isotope. The authors thank V. I. Baranov  
for his interest in this work. There are 6 figures and 2 references:  
1 Soviet and 1 Canadian.

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S/020/60/134/006/029/031  
B004/B054

AUTHORS: Lavrukhina, A. K., Kolesov, G. M., and Tan Syao-yen

TITLE: Reduction of Rare Earths of the Cerium Group on the Mercury Cathode

PERIODICAL: Doklady Akademii nauk SSSR, 1960, Vol. 134, No. 6, pp. 1406-1409

TEXT: The authors wanted to investigate the behavior of samarium, promethium, and berium in electrolysis. Mercury was used as cathode, platinum wire as anode. The electrolysis was carried out at pH 5.8 - 6.0, at 0 to +2°C, a concentration ratio of  $\text{Me}^{3+} : \text{CH}_3\text{COO}^- : \text{C}_6\text{H}_5\text{O}_3^- = 1 : 1 : 2$ , and a voltage of 10 - 12 v. Radiochemically pure  $\text{Sm}^{153}$ ,  $\text{Pm}^{147}$ , and  $\text{Ce}^{144}$  were used. After electrolysis, the quantity of Sm, Pm, and Ce, which had passed over into Hg, was determined by an MCT-17 (MST-17) counter. Sm was also determined gravimetrically. The experimental data are shown in Table 1 and Figs. 1, 2. The separation of samarium depends on the current density only in the range between 25 and 75 ma/cm<sup>2</sup>. At higher current density,

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Reduction of Rare Earths of the Cerium Group  
on the Mercury Cathode

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the separation is little influenced by it. At  $100 \text{ ma/cm}^2$ , already 95% of Sm are separated within 30 min. At a constant concentration of the complex formers, the Sm separation does not depend on them within concentrations from  $6 \cdot 10^{-3}$  to  $6 \cdot 10^{-2}$ .  $\text{Pm}^{147}$  and  $\text{Ce}^{144}$  were used in concentrations of about  $1 \cdot 10^{-9}$  and  $1 \cdot 10^{-12}$  mole. Pm passes over into the amalgam only at current densities higher than  $75 \text{ ma/cm}^2$ . But even at  $100 \text{ ma/cm}^2$ , the Pm separation only starts after 30 min and after the formation of a samarium-potassium amalgam with corresponding potential ( $-2.11 \text{ v}$ ). The Pm separation depends on the alkali metal used (K, Li). In the presence of K, the Pm only passes over into the Hg cathode if large amounts of Sm or Yb are present. In the presence of Li, the Pm separation is independent of the presence of these rare earths. An addition of potassium citrate bears no influence on the Sm separation, but prevents that of Pm, whereas lithium citrate exerts no influence on the Pm separation. The same results were obtained for cerium as for promethium. The authors arrived at the conclusion that the passing over of Sm, Pm, and Ce into the Hg cathode is due to electrolytic reduction. An intermediate stage is the bivalent state which is polarographically confirmed for La, Ce, Pr, and Nd (Refs. 8-10):

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Reduction of Rare Earths of the Cerium Group  
on the Mercury Cathode

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$\text{Me}^{3+} + e^- \longrightarrow \text{Me}^{2+}$ . The authors found the optimum conditions for the quantitative separation of Sm (up to 99.7% within an hour), and established that Pm and Ce pass over into the Hg cathode up to 97%, even if the complex formers are present in a  $10^{11}$  excess. They also clarified the possibility of a separation of these elements. They thank S. I. Sinyakova, Yu. S. Sklyarenko, and O. L. Kabanova for their discussion. There are 2 figures, 1 table, and 10 references: 7 Soviet, 2 US, and 1 German.

ASSOCIATION: Institut geokhimii i analiticheskoy khimii im. V. I. Vernadskogo  
(Institute of Geochemistry and Analytical Chemistry imeni V. I. Vernadskiy)

PRESENTED: June 8, 1960, by A. P. Vinogradov, Academician  
SUBMITTED: May 30, 1960

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LAVRUKHINA, A.K.

Report to be submitted for the IUPAC 21st Conference and 18th Intl. Congress of Pure and Applied Chemistry, Montreal, Canada, 2-12 August 1961

GORUNITSKY, G. V., Academy of Sciences USSR, Kiev - "The oscillographic investigation of the electrochemical kinetics in fused salts" (Section A.3, s.2 - Session I, 11 Aug 61, afternoon)  
 GURVICH, I. V., Academy of Sciences USSR, Moscow - "The calculation of thermodynamic functions of gases in a wide temperature range" (Section A.3, c.(1), Session II - 8 Aug 61, afternoon)  
 KAPIN, V. A., Physico-Chemical Institute Ireni I. Ya. Karpov, Moscow - "Vitrification phenomena in crystalline polymers" (Section B.4 - 7 Aug 61, afternoon)  
 KIMURA, A. F., Moscow State University Ireni M. V. Lomonosov - "The influence of surface heterogeneity and adsorbate-adsorbate interaction on the adsorption properties of solid surfaces" (Joint Session, Sections A.2 and B.1 - 8 Aug 61, morning)  
 KUDAS'KOV, V. B., Institute of Chemical Physics, Academy of Sciences USSR, Moscow - "The H<sub>2</sub>O radical" (Section A.1, Session I - 8 Aug 61, morning) (Also, Section A.1, Chairman, Session I - 8 Aug 61, morning) (Also, Section A.1, Chairman, Session I - 8 Aug 61, morning)  
 KURCHOV, V. I., Institute of Geochemistry and Analytical Chemistry Ireni V. I. Vernadsky, Academy of Sciences USSR - "A novelty in the use of organic compounds for concentration of small amounts of the elements" (To be presented in Russian) (Section C.2 - 11 Aug 61, morning)  
 KURCHUK, A. K., RANVICH, L. Z., and IYOSKALOVA, L. P., Institute of Geochemistry and Analytical Chemistry Ireni V. I. Vernadsky, Academy of Sciences USSR - "New data on radiolchemical investigations of the processes of fission and fragmentation induced by high energy protons" (Section A.4 - 8 Aug 61, afternoon)  
 LOMONOSOV, M. V., Academy of Sciences USSR, Moscow - "Determination of rate constants of elementary processes in chemical reactions as a function of temperature, pressure, and molecular transfer coefficients" (Section A.3, b, (2) - 7 Aug 61, afternoon)  
 MAMONTEV, S. (Probably MUDRACH, S.) and GERASIMOV, I. I., Moscow State University Ireni M. V. Lomonosov - "Study of the thermodynamic properties of the iron-iron-lanthanum" (Section A.3, c.(3), Session II(A) - 11 Aug 61, morning)  
 PAVLOV, G. M., MOKHOMEDOV, A. M., MALAYEV, V. F., and SHEN, I. A., Moscow State University Ireni M. V. Lomonosov - "Smelting of complex ores in solid-phase reactions" (Joint Session, Sections A.2 and B.1, 8 Aug 61, morning)  
 SPATSKY, H., Institute of Chemical Physics, Academy of Sciences USSR, Moscow - "Chemical reactions at reduced temperatures and related problems of energy transfer" (To be presented in Russian) (Plenary lecture - Saturday, 12 Aug 61)  
 SMILG, I. A., Academy of Sciences USSR, Moscow - "The active ligands and the intermolecular complexes in the tetrahedral complexes of halogenation of the organic compounds" (Section A.1, Session II - 11 Aug 61, afternoon)  
 SMIRNOV, M. I., Electrochemistry Institute, Dnepropetrovsk - "The equilibrium between the titanium subgroup metals and the salt melts" (Section B.3 - 7 Aug 61, afternoon)  
 TAYLOR, V. L., Institute of Chemical Physics, Academy of Sciences USSR - "Reaction of ions and molecules in the gas phase" (Section A.1, Session I - 9 Aug 61, afternoon)  
 TROTSKY, Alexander N., Leningrad State University Ireni A. A. Zhdanov - (Section A.1, Session I - 8 Aug 61, afternoon Session) (Also on program for Section A.1, Session I - 8 Aug 61, afternoon)  
 TROTSKY, Alexander N., Leningrad State University Ireni A. A. Zhdanov, N. V. Leningrad State University Ireni A. A. Zhdanov, N. V. Leningrad State University Ireni A. A. Zhdanov - "Mass-spectrometry and fluorescence of radicals in the photolysis of organic compounds" (Section I.9 Aug 61, afternoon)  
 ULTRA-VIOLET RADIATION" (Section A.1, Session I - 9 Aug 61, afternoon)  
 YAKOVLEV, M. B., Scientific Research Physico-Chemical Institute Ireni I. Ya. Karpov - "On the dissociation of molecules on electron impact and the early stages of radiation-chemical processes" (Section A.1, Session I - 8 Aug 61, afternoon)  
 YAKOVLEV, M. B., and MOSEV, V. V., Institute of Geochemistry and Analytical Chemistry Ireni V. I. Vernadsky, Moscow - "The plasma generator and its use for spectral analysis of alloys and rocks" (Section C.1 - 8 Aug 61, morning)  
 YAKOVLEV, A. F., LOMONOSOV, A. K., and REYKA, L. D., Institute of Geochemistry and Analytical Chemistry Ireni V. I. Vernadsky, Academy of Sciences USSR - "The study of nuclear reactions in iron meteorites under the action of high energy protons" (Section A.4 - 8 Aug 61, afternoon)  
 YAKOVLEV, M. B., and ALDURIN, I. I., Institute of Geochemistry and Analytical Chemistry Ireni V. I. Vernadsky, Academy of Sciences USSR - "The determination of trace impurities in some materials for activation analysis" (Section C.1 - 8 Aug 61, afternoon)  
 YAKOVLEV, Boris V., Institute of Physical-Chemical Chemistry, Minsk - "The effect of donor and acceptor substituents on the decomposition rate of solids" (Section A.2 - 8 Aug 61, afternoon)

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29394  
S/007/61/000/011/001/003  
B107/B147

AUTHORS: Vinogradov, A. P., Lavrukhina, A. K., Revina, L. D.

TITLE: Nuclear reactions in iron meteorites

PERIODICAL: Geokhimiya, no. 11, 1961, 955 - 966

TEXT: The authors report on a radiochemical analysis of the fission products of iron bombarded with 660-Mev protons. They attempted to clarify the cosmogenic formation of various isotopes in iron meteorites. The synchrocyclotron of the Laboratoriya yadernykh problem Ob'yedinennogo instituta yadernykh issledovaniy (Laboratory for Nuclear Problems of the Joint Institute of Nuclear Research) was used to bombard 100 to 500 mg of iron powder with about  $10^{12}$  protons/sec·cm<sup>2</sup> for 0.5 to 2 hr. The resulting isotopes were identified according to half-life, kind and energy of radiation. A simplified magnetic beta spectrometer and a gamma scintillation spectrometer were used for this purpose. A total of 38 isotopes with atomic numbers 4 - 27 and half-lives from 8 min to 3 years were found. The production cross sections and yields of stable and undetected radioisotopes were calculated by interpolation (Fig. 2). On the strength of

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# Nuclear reactions in iron meteorites

these data, the number of cosmogenic nuclei was calculated, which are formed by fission of  $\text{Fe}^{56}$  in the center of a meteorite of 10 cm diameter within  $4.5 \cdot 10^9$  years (Table 4). Results: Within the period mentioned, about  $10^{-7}$  g of cosmogenic isotopes per gram of meteorite is formed, among them the stable isotopes  $\text{Ar}^{36}$ ,  $\text{Ar}^{38}$ ,  $\text{K}^{40}$ ,  $\text{Sc}^{45}$ , and  $\text{V}^{50}$ . The concentration calculated for these isotopes agrees with the observed concentration and is about  $10^{-9}$  g/g of meteorite. This explains the anomalies observed in the isotopic composition of potassium and argon. Shifts toward the ratios in terrestrial rocks are to be expected also for the isotopes of vanadium, titanium, and other elements. The equilibrium values for the activity of long-lived cosmogenic nuclei in iron meteorites were calculated. A comparison with values measured in various meteorites shows deviations by a factor of  $\leq 5$ . The mean production cross section for nuclei with an atomic weight of about 20 and about 40 was found to be 1.2 and 0.5 Bev. The authors thank V. V. Malyshev, L. M. Saratova, and Su Hung-kuei for help in the experimental work. L. K. Levskiy and V. Kuznetsov are mentioned. There are 4 figures, 7 tables, and 30 references. Card 2/5

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S/186/61/003/006/006/010  
E051/E135

AUTHORS: Lavrukhina, A.K., Kolesov, G.M., and T'ang Hsiao En

TITLE: The separation of irradiated samarium from  
radioactive rare earth elements by electrolysis  
with a mercury cathode

PERIODICAL: Radiokhimiya, v.3, no.6, 1961, 724-731

TEXT: Neutron-deficient elements of the cerium group of rare earth elements can be produced by bombardment, with high-energy particles, of a samarium oxide target weighing 100 mg or more. After the irradiation the major part of the samarium must be removed before using an ion-exchange chromatographic procedure for the complete separation of the isotopes produced. Previous experience had shown that electrolysis with a mercury cathode would serve to remove most of the samarium, although other rare earth elements, especially promethium, showed a tendency to be co-deposited. Various factors were studied in an attempt to improve the samarium removal and decrease the other rare earth loss. Electrolysis was carried out in a cell made from a  
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separating funnel (37 mm diameter), ice-jacketed to keep the cell at 0-2 °C. The cathode was 15 ml of purified mercury and the anode 25 cm of 0.8-1.0 mm diameter platinum wire, wound in a spiral. Current to the cell was supplied at 10-12 volts from a stabilised selenium rectifier circuit. Both mercury and electrolyte were stirred at 500 r.p.m. The samarium target material was shown by spectrographic analysis to contain ~0.5% Eu<sub>2</sub>O<sub>3</sub>, < 0.01% Gd<sub>2</sub>O<sub>3</sub>, and small amounts of Cu, Fe, Al, Si, Mg and Mn. Radioactive tracers Sm<sup>153</sup>, Pm<sup>147</sup> and Ce<sup>144</sup> were added as required. The radiochemical purity of the tracers was checked by half-life measurements, β-ray absorption in aluminium and by the absence of γ-radiation. The solution for electrolysis was prepared by dissolving the target material in 10 ml of 25% acetic acid, excess of which was removed by evaporating the solution until crystals of samarium acetate appeared. The residue was dissolved in 6-8 ml of 3.8% potassium citrate solution and transferred to the cell using 2-4 ml of the potassium citrate solution as washings, giving a solution of pH 5.8-6.0. After an electrolysis the cathode mercury was run out of the cell with

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the current switched on still, and washed with water to decompose potassium amalgam. The remaining samarium amalgam was decomposed with 6N HCl and samarium oxalate precipitated. From the exhausted electrolyte  $\text{SmF}_3$  was precipitated; this was dissolved in 30%  $\text{HNO}_3$  saturated with  $\text{H}_3\text{BO}_3$  and  $\text{Sm}(\text{OH})_3$  or  $\text{Sm}_2\text{O}_3$  precipitated. The precipitates were ignited at 85 °C and weighed. 0.05-0.10 m aliquots of the electrolyte were assayed for  $\text{Sm}^{153}$ ,  $\text{Pm}^{147}$  and  $\text{Ce}^{144}$  activity. A loss of about 1% of the activity occurred due to absorption on the glass walls of the cell. 0.1-0.3% loss occurred when the potassium amalgam was decomposed, and other losses amounted to about 0.02%. The effects of electrolysis time, cathodic current density, samarium concentration and the substitution of lithium citrate for potassium citrate were examined, keeping the citrate concentration constant. At 100 mamp/cm<sup>2</sup> current density, after 30 min. electrolysis, 4.6% of the samarium and 99.9% of the promethium were left in the electrolyte. After 45 min the figures were 3.6% samarium and 59% promethium remaining in the electrolyte. Cerium was removed from solution slowly but continuously, the deposition increasing

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with increasing cerium concentration. At current densities of 25 and 50 mamp/cm<sup>2</sup> no promethium was deposited in 1.5 hours, but at 100 mamp/cm<sup>2</sup> 83% was deposited. Samarium deposition increased sharply from 25 to 75 mamp/cm<sup>2</sup> with a slight increase at higher current densities. From 10<sup>-6</sup>-10<sup>-4</sup> M samarium the quantity deposited remained ~93%; from 10<sup>-4</sup>-6 x 10<sup>-2</sup> M the quantity deposited rose to 97.6%; above 6 x 10<sup>-2</sup> M the quantity decreased due to precipitation of basic samarium acetate. By replacing potassium citrate by lithium citrate, the deposition of samarium was improved, leaving 0.3% in the electrolyte after one hour's electrolysis at 100 mamp/cm<sup>2</sup>, but at the same time 96.8% of the promethium was deposited in the mercury cathode. The optimum conditions for the separation of samarium on a mercury cathode in the presence of potassium citrate were found to be:

6 x 10<sup>-2</sup> to 6 x 10<sup>-3</sup> M samarium concentration, 100 mamp/cm<sup>2</sup> current density, 60 minutes electrolysis time, and molar ratio Sm<sup>3+</sup>:Cit<sup>3-</sup> = (1:2) - (1:20). Under these conditions the yield of

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samarium reaches 97.6%. By replacing K with Li, a yield of 99.7% was obtained. Previous work had shown that promethium and cerium only start to be deposited in a mercury cathode after the formation of a mixed potassium-samarium amalgam, and the extent of deposition depends on the potassium citrate concentration. By changing the cathode mercury during the course of an electrolysis an improvement is gained in samarium deposition without any deposition of promethium. With two changes of mercury at 100 mamp/cm<sup>2</sup>, 97.5% of the samarium can be removed from the electrolyte without loss of promethium. With three changes of mercury, 9% of the promethium was deposited. If 400 mg of potassium citrate was added during the course of an electrolysis, although no improvement in samarium separation occurred, no promethium was removed from the electrolyte at all. Using these latter conditions a samarium oxide target which had been bombarded with 660 MeV protons was treated. After the major part of the samarium had been removed by electrolysis an ion-exchange chromatographic procedure separated isotopes of Sm, Pm, Nd, Pr, Ce and La.

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The separation of irradiated ...

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There are 5 figures, 5 tables and 19 references; 13 Soviet-bloc, 1 Russian translation from non-Soviet-bloc publication, and 5 non-Soviet-bloc. The English language references read as follows;

- Ref. 2: H.N. McCoy, J. Am. Chem. Soc., v. 63, 6, 1622 (1941).
- Ref. 3: H.N. McCoy, J. Am. Chem. Soc., v. 63, 12, 3432 (1941).
- Ref. 4: I.K. Marsh, J. Chem. Soc., 531 (1943).
- Ref. 11: E.I. Onstott, J. Am. Chem. Soc., v. 78, 10, 2070 (1956).

SUBMITTED: May 26, 1960

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S/056/61/040/002/003/047  
B113/B214

AUTHORS: Lavrukhina, A. K., Rakovskiy, E. Ye., Su Khun-guy,  
Khoynatskiy, S.

TITLE: Nuclear fission of antimony by high-energy protons

PERIODICAL: Zhurnal eksperimental'noy i teoreticheskoy fiziki, v. 40,  
no. 2, 1961, 409-418

TEXT: The nuclear fission products of antimony due to 660-Mev protons have been investigated to obtain the main characteristics of this process: mass spectrum, isotopic composition of the fission fragments, distribution of the nuclear charge, and the amount of the cross section. The target for irradiation was prepared from metallic antimony which was purified (spectrally pure) by repeated zone melting. The antimony target was coated with aluminum which served as a monitor for the determination

of the proton flux according to the reaction  $Al^{27}(p, 3pn)Na^{24}$ . This target was irradiated in the inner beam of the synchrocyclotron of the LYaP OIYaI (Laboratory for Nuclear Problems of the Joint Institute of Nuclear Research) for 0.5 - 3 hr. The elements of atomic numbers 11-37

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Nuclear fission of antimony...

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were separated by chemical methods. The activity of the preparations was determined by the MCT-17 (MST-17) end-window counter. For the identification of the individual activities, the sign of the radiation was determined in a magnetic analyzer. Two groups of products could be identified from the data on half-life, mode of disintegration, and mean production cross section. Between Rb and Zn ( $Z \geq 30$ ) are isotopes whose yield decreases rapidly with increasing  $\Delta Z = Z_0 - Z$  ( $Z_0$  is the atomic number of the initial nucleus). The range  $16 \leq Z \leq 28$  is to be attributed to the fission products for which no change in the yield was observed with a change in  $Z$ . The interpolation method was used for estimating the yield of the unidentified, stable, long- and short-lived radioactive fission fragments from antimony. It is seen from Fig.2 that the main part of fission fragments lies in the immediate neighborhood of the broken line of stable nuclei. The character of distribution of the fragment yield from antimony in  $A$  and  $Z$  can be determined from the totality of the experimental and interpolated data. The mass distribution curve of the isotopic yield is dome-shaped. On the fission

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Nuclear fission of antimony...

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B113/B214

of antimony nucleus there occurs a considerable increase in the relative amount of neutron-deficient isotopes.  $\sigma$  and the amount of the most probable charge  $Z_p$  was determined for all isobars between  $A = 37$  and  $A = 62$ . It was found that  $Z_p(A)$  lies very near the line of stable nuclei. The charge distribution is constant for all the isobars. Fig. 8 shows the distribution curves of the total yield as a function of  $Z$  for the fission fragments of antimony, holmium, bismuth, and uranium nuclei. It is also seen from Fig. 8 that the increase in the nuclear charge of the target makes the curve broader which indicates that the contribution of the asymmetric fission increases with increasing charge of the fissioned nucleus. A comparison of the curves in Fig. 8 shows that the yield of fission fragments of antimony nuclei is significantly smaller than that of the heavier nuclei. The total cross section for nuclear fission of antimony by 660-Mev protons was found to be 0.25 mb. From the results obtained it is possible to conclude that a regular change in all the fundamental parameters of the fission process is connected with the change in the charge of the target nucleus. V. N. Mekhedov and T. B. Malysheva are thanked for their valuable advice; L. D. Revina,

Card 3/7

Nuclear fission of antimony...

S/056/61/040/002/003/047  
B113/B214

L. D. Firsova, and I. S. Kalicheva are thanked for their help in the experimental part of the work. Yu. V. Yakovlev, L. A. Smakhtin, V. Shamov, and V. V. Malyshev are mentioned. There are 8 figures, 2 tables, and 18 references: 15 Soviet-bloc and 3 non-Soviet-bloc.

ASSOCIATION: Institut geokhimii i analiticheskoy khimii Akademii nauk SSSR (Institute of Geochemistry and Analytical Chemistry of the Academy of Sciences USSR)

SUBMITTED: July 8, 1960

Legend to Table 1: Yields of identified fission fragments from antimony bombarded by 660-Mev protons. 1) Element, 2) atomic weight, 3) mode of disintegration, 4) experimental half-life, 5) half-life taken from tables. ЭЗ - electron capture, ИП - isomeric transition, час - hour, дн - day, мин - minute.

Card 4/7

S/020/61/137/003/008/030  
B104/B214

AUTHORS: Lavrukhina, A. K., Malysheva, T. V., and Khotin, B. A.

TITLE: A new Ir<sup>183</sup> isotope

PERIODICAL: Doklady Akademii nauk SSSR, v. 137, no. 3, 1961, 551-552

TEXT: On separation of the daughter osmium from the iridium fraction produced by the disintegration of gold by 660-Mev protons, an activity with the half-life of 12 hours was discovered which belongs to Os<sup>183</sup>. This fact confirms the formation of a new neutron deficient iridium isotope with the mass number 183 in the disintegration product of gold. The present paper is concerned with the determination of the half-life of Ir<sup>183</sup>. 0.5 g of metallic gold was irradiated for 0.5 - 1 hr by 660-Mev protons with the synchrocyclotron of the Ob'yedinennyy institut yadernykh issledovaniy (Joint Institute of Nuclear Research). The radioactive, pure iridium was obtained from the disintegration products of gold by a method described in an earlier paper of the present authors (Ref. 3: A. K. Lavrukhina et al. III soveshch. po neytronodefitsitnym

Card 1/4



A new Ir<sup>183</sup> isotope

S/020/61/137/003/008/030  
B104/B214

izotopam, g. Dubna, 1960). Then, the daughter osmium was separated from the iridium in the same time interval. (The time of accumulation was changed in the different experiments from 1.5 to 4 hr). For this purpose, 20 mg of osmium in the form of Na<sub>2</sub>OsO<sub>4</sub> was added to a solution of the radioactive iridium. The OsO<sub>4</sub> was extracted with hydrochloric acid and fixed with 10% NaOH. Thereafter the osmium sulfide was precipitated and annealed to metal in a current of hydrogen. The degree of separation of osmium was checked by weighing the metallic osmium. The chemical yield was 95-98%. The decrease in radioactivity was measured by an end-window counter. Fig. 1 shows the decay curves of the activity of osmium determined at intervals of 1.5 hr for the 4th and 7th separations. Half-lives of 10 minutes, 12 hr and 90 days were measured. It is obvious that the activity caused by Os<sup>183</sup> is the greatest part of the activity of osmium. In the osmium separated 12 hr after the first iridium separation (9th separation, Fig. 2) the 12-hr activity was no more present. From Fig. 3 appears a half-life of  $1 \pm 0.1$  hr for Ir<sup>183</sup>. An activity with the half-life  $1 \pm 0.15$  hr could be established also from the decay curve of the total activity of the iridium fraction. The activity with

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A new Ir<sup>183</sup> isotope

S/020/61/137/003/008/030  
B104/B214

the half-life of 90 days is due to the formation of Os<sup>185</sup> from the 15-hr Ir<sup>185</sup>. The activity with the half-life of 10±2 minutes belongs to the osmium isomer Os<sup>190m</sup>. There are 3 figures and 3 Soviet-bloc references.

ASSOCIATION: Institut geokhimii i analiticheskoy khimii im. V. I. Vernadskogo Akademii nauk SSSR  
(Institute of Geochemistry and Analytical Chemistry imeni V. I. Vernadskiy, Academy of Sciences USSR)

PRESENTED: October 10, 1960, by A. P. Vinogradov, Academician

SUBMITTED: September 23, 1960

Card 3/4

21489

S/020/61/137/004/013/031  
B104/B206

24.6600 (1138, 1098)  
AUTHORS: Lavrukina, A. K., Rakovskiy, E. Ye., Su Hung-kuei, and  
Khoynatskiy, S.

TITLE: Fast-proton induced fission of antimony nuclei

PERIODICAL: Doklady Akademii nauk SSSR, v. 137, no. 4, 1961, 826-829

TEXT: The difficulties in the experiments described here mainly consisted in that the fission fragments of antimony nuclei possess very small yields as compared with those of the disintegration products.  $10^{-3}$  % of impurities lead to strong deviations from the correct results. High-purity targets had therefore to be used. The targets were irradiated in the synchro-cyclotron of the Ob'yedinennyy institut yadernykh issledovaniy (Joint Institute of Nuclear Research) with 660 Mev protons; the products were chemically separated ( $Z = 11 - 37$ ). Great difficulties occurred here too, since antimony fission products are often identical with antimony disintegration products. The isotope production cross sections were calculated by conventional methods. The results are compiled in Table 1. As can be seen from the diagram in Fig. 1, the fission of Sb takes place in a much

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S/020/61/137/004/013/031  
B104/B206

Fast-proton induced fission of...

wider interval of the ratio  $n/p$  as is the case for heavy nuclei. With it, however, the share of neutron-deficient nuclei is also greater than for heavy nuclei. Most of the nuclei identified by the authors are "protected" (zashchishchenny) isobars, which makes it possible to determine the little known distribution of the nuclear charge on the fission of Sb. The distribution of the isobaric output is shown in Fig. 2. The half-widths of the curves for the individual isobars are 3-4 unit charges, while the same half-widths amount to 2-3 unit charges for the fission of heavy nuclei. The line which connects the most probable nuclear charges of the fragments lies close to the line of nuclear stability. The fact is also mentioned that the Sb fission takes place symmetrically (Fig. 2), which is similar to the fission of Ag. With a reduction of the proton energies to 220 Mev, the portion of asymmetric fission products is reduced. As it turned out, the fission of Sb is accompanied by an average emission of 7 protons. From Table 2 it can be seen that the fission cross section increases with increasing  $Z$  of the target nucleus. The total fission cross section for Sb with 660 Mev-protons is 0.25 millibarn. This value almost equals that determined on Ag for the same proton energies. (0.32 millibarn). The authors thank V. N. Mekhedov, L. D. Revina and L. P. Moskaleva for advice

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S/020/61/137/004/013/031  
B104/B206

Fast-proton induced fission of...

and assistance. There are 3 figures, 2 tables, and 13 references:  
9 Soviet-bloc and 4 non-Soviet-bloc.

ASSOCIATION: Institut geokhimii i analiticheskoy khimii im. V.I. Vernadskogo  
Akademii nauk SSSR (Institute of Geochemistry and Analytical  
Chemistry imeni V.I. Vernadskiy of the Academy of Sciences  
USSR)

PRESENTED: October 10, 1960, by A. P. Vinogradov, Academician

SUBMITTED: September 23, 1960.

Card 3/7

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S/020/61/140/004/015/023

B106/B110

AUTHORS: Kourzhim, V., Lavrukhina, A. K., and Rodin, S. S.

TITLE: Use of ammonium phosphotungstate for the separation of rubidium and cesium by ion exchange

PERIODICAL: Akademiya nauk SSSR. Doklady, v. 140, no. 4, 1961, 832-834

TEXT: J. van R. Smith and co-workers (Ref. 10, see below) recently published a paper on the separation of alkali metals on an exchanger column with ammonium phosphomolybdate. J. Krtil and V. Kourim (Ref. 11, see below) showed that the chemical stability of ammonium phosphotungstate in neutral and highly acid solution was higher than that of ammonium phosphomolybdate, and that the sorption capacity of the former salt was twice that of the latter. On the basis of these data, the authors studied the separation of the heavy alkali metals rubidium and cesium on a column filled with ammonium phosphotungstate. To reduce the high hydraulic resistance of fine-crystalline ammonium phosphotungstate, a coarse packing had to be added. A fine-fibered tremolite asbestos of the amphibolic type was used for this purpose. Silica gel, glass wool.

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29014

S/020/61/140/004/015/023

B106/B110

Use of ammonium phosphotungstate ...

cellulose, and cork crumbs were less suitable. The column used had an inside diameter of 5 mm, and was filled with a suspension of asbestos in 1 M  $\text{NH}_4\text{NO}_3$  solution up to a height of about 30 mm. A 0.2 M solution of phosphotungstic acid, and then a 0.3 M  $\text{NH}_4\text{NO}_3$  solution, were passed through this column. 0.1 ml of a 1 M  $\text{HNO}_3$  solution was then introduced which contained  $10^{-4}$  M rubidium (radiolabeled with  $\text{Rb}^{86}$ ) and  $10^{-6}$  M cesium (radiolabeled with  $\text{Cs}^{134}$ ). The absorption of  $\beta$ -radiation by an aluminum filter was used to identify the activities since the energies of  $\beta$ -particles of  $\text{Rb}^{86}$  (1.79 Mev) and  $\text{Cs}^{134}$  (0.65 Mev) are highly different. A filter of a thickness of 204 mg/cm<sup>2</sup> was used. The radioactivity of fractions obtained after separation on the column was measured in an end-window counter of the MCT-17 (MST-17) type with and without aluminum filter. The relative quantities of the two active isotopes in the fractions were calculated from the following equations:  

$$A_{\text{Rb}} = 6.82 \cdot A_{\text{Al}} - 0.31 \cdot A_{\text{tot}}; A_{\text{Cs}} = 1.31 \cdot A_{\text{tot}} - 6.82 \cdot A_{\text{Al}}; (A_{\text{Rb}}, A_{\text{Cs}} - \text{activities of } \text{Rb}^{86} \text{ and } \text{Cs}^{134}, \text{ respectively; } A_{\text{Al}} - \text{total activity when})$$

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Use of ammonium phosphotungstate ...

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S/020/61/140/004/015/023  
B106/B110

measuring with filter;  $A_{tot}$  - total activity without filter). Rubidium was eluted from the column, with 1 M ammonium nitrate solution cesium with 6 M ammonium nitrate solution. Fig. 1 shows the resultant chromatogram. Rb and Cs can also be separated by ammonium silicomolybdate, but this salt is more soluble in  $NH_4NO_3$  solution than ammonium phosphotungstate, and is therefore, slowly eluted from the column. The chromatograms obtained with ammonium phosphomolybdate agree with data in Ref. 10 (see below). Ammonium silicotungstate cannot be applied to chromatographic separations because of its good solubility in ammonium nitrate solutions. The chromatographic separation method described may be valuable for the separation of highly active isotopes of rubidium and cesium since phosphotungstates are very stable to radiation. There are 1 figure and 16 references: 3 Soviet and 13 non-Soviet. The three most recent references to English-language publications read as follows: Ref. 10: J. van R. Smith, W. Robb, I. I. Jacobs, J. Inorg. and Nucl. Chem., 12, 104 (1960); Ref. 11: J. Krtil, V. Kouřim, J. Inorg. and Nucl. Chem., 12, 367 (1960); A. K. Lavrukchina, A. A. Pozdnjakov, S. S. Rodin, Intern. J. of Appl. Rad. and Isotopes, 2, No 1-4, 34 (1960).

Card 3/4



Use of ammonium phosphotungstate ...

29014

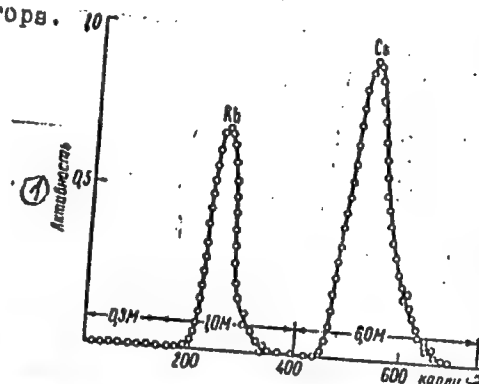
S/020/61/140/004/015/023  
B106/B110

ASSOCIATION: Institut geokhimii i analiticheskoy khimii im. V. I. Vernadskogo Akademii nauk SSSR (Institute of Geochemistry and Analytical Chemistry imeni V. I. Vernadskiy of the Academy of Sciences USSR)

PRESENTED: April 6, 1961, by A. P. Vinogradov, Academician

SUBMITTED: March 4, 1961

Legend to Fig. 1: (1) activity; (2) drops.



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LAVRUKHINA, A. K.

JUN 25 1963

RUN/6244

PHASE I BOOK EXPLOITATION

Lavruhina [Lavrukhina], A. K.

Succesele chimiei nucleare. Bucharest, Ed. Stiintifica, 1962. 144 p.  
Errata slip inserted. 5100 copies printed.

Transl. of Uspekhi yadernoy khimii (Achievements in Nuclear Chemistry).  
Moskva, Izd-vo AN SSSR, 1959. 143 p. (Series: Akademiya nauk  
SSSR. Nauchno-populyarnaya seriya)

Ed.: P. Hodorogea; Tech. Ed.: Gh. Popovici.

PURPOSE: This book is intended for general readers interested in  
nuclear chemistry.

COVERAGE: The book, a translation from the Russian, is a review of  
the progress made in nuclear chemistry up to 1962.

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Achievements in Nuclear Chemistry

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Conclusion

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Bibliography

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AVAILABLE: Library of Congress

SUBJECT: Chemistry

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BN/wro/jk  
5/3/63

LAVRUKHINA, Avgusta Konstinovna; KOLESOV, Gennadiy Mikhaylovich;  
PODUSHVINA, V.A., red.; MAZEL', Ye.I., tekhn. red.

[Formation of chemical elements in cosmic bodies]Obrazovanie  
khimicheskikh elementov v kosmicheskikh telakh. Moskva, Gos-  
atomizdat, 1962. 171 p. (MIRA 15:12)  
(Chemical elements) (Cosmogony)

IAVRUKHINA, A.

Stars narrate. Nauka i zhyttia 12 no.6:20-22, 1 of cover Je '62.  
(MIRA 15:7)

(Chemical elements) (Stars)

S/056/62/043/001/001/056  
B154/B108

AUTHORS: Lavrukhina, A. K., Moskaleva, L. P., Malyshev, V. V.,  
Safarova, L. M.

TITLE: Production of light nuclei by bombarding heavy elements with  
660 Mev protons

PERIODICAL: Zhurnal eksperimental'noy i teoreticheskoy fiziki, v. 43,  
no. 1(7), 1962, 3-7

TEXT: The authors investigate the cross sections  $\sigma$  for the production of  
Be<sup>7</sup>, F<sup>18</sup>, Na<sup>24</sup>, Mg<sup>28</sup>, Si<sup>31</sup>, P<sup>32</sup> by 660 Mev proton bombardment of Al, Cu,  
Sb, Sn, Bi, U. The relative contributions of fission and fragmentation  
in Na<sup>24</sup> production are estimated from the energy and angular distributions  
of the Na<sup>24</sup> nuclei produced by bombarding Cu. The Al, Cu, Sb, and U targets  
were bombarded in the usual way (A. K. Lavrukhina, et al. Atomn. energ.,  
3, 285, 1957); Sn and Bi were kept in special graphite containers.  
The authors conclude that the production of Si<sup>31</sup> and P<sup>32</sup> by bombarding

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S/056/62/043/001/001/056  
B154/B108

Production of light nuclei by ...

Cu and neighboring elements is a result of spallation and symmetric fission. Formation of lighter isotopes from all target nuclei occurs via

fission and fragmentation. The ratio  $\frac{\sigma(\text{Na}^{24})}{\sigma(\text{F}^{18})}$  is always  $> 1$  and amounts

to 2.5, 5.0, 2.8, 1.3 and 1.8 for Cu, Sb, U, Bi and Sn, respectively.

The measured values of  $\sigma$  in the bombardment of Bi are virtually equal for all light nuclei which may be due to the spherical symmetry of these nuclei.

The energies of the fragments from Cu fission ( $\text{Na}^{24}$  nuclei) in the angular interval of  $15-80^\circ$  are greater and the energies in the angular interval

of  $100-160^\circ$  are smaller than the Coulomb repulsion of  $\text{Na}^{24}$  (20 Mev) so that asymmetric fission is supposed. The considerable anisotropy observed in the angular interval of  $10-30^\circ$  and the fragments with energies greater than that of Coulomb repulsion are indicative of fragmentation contributing to the process. The integral yield in fragments of a certain type depends on the "separation energy"  $E = m_B + m_F - m_A$  ( $m_A$  - mass of target

nucleus,  $m_F$  - mass of fragment,  $m_B$  - mass of additional fragment).

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Production of light nuclei by ...

S/056/62/043/001/001/056  
B154/B108

There are 2 figures and 3 tables.

ASSOCIATION: Institut geokhimii i analiticheskoy khimii Akademii nauk SSSR  
(Institute of Geochemistry and Analytical Chemistry of the  
Academy of Sciences USSR)

SUBMITTED: December 26, 1961 (initially)  
March 27, 1962 (after revision)

Card 3/3

TRIFANOV, Dmitriy Nikolayevich; LAVRUKHINA, A.K., doktor khim.  
nauk, otv. red.; ALMAZOV, A.B., red. izd-va; DOROKHINA,  
I.N., tekhn. red.

[If there were no uranium and thorium] Esli by ne bylo urana  
i toriia. Moskva, Izd-vo Akad. nauk SSSR, 1963. 85 p.  
(MIRA 16:5)

(Chemical elements) (Nuclear geophysics)

LAVRUKHINA, Avgusta Konstantinovna; MALYSHEVA, Tamara Vladimirovna;  
PAVLOTSKAYA, Fanni Il'ichna; BARANOV, V.I., prof., otv.  
red.; DRAGUNOV, E.S., red.; GUSEVA, A.P., tekhn. red.

[Radiochemical analysis] Radiokhimicheskii analiz. Moskva,  
Izd-vo AN SSSR, 1963. 219 p. (MIRA 16:12)  
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LAVRUKHINA, A.K.; YUKINA, L.V.; KHROMCHENKO, Z.V.

Extraction of rare-earth elements. Trudy Kom.anal.khim. 14:  
202-208 '63. (MIRA 16:11)

LAVRUKHINA, A.K.; REVINA, L.D.; MALYSHEV, V.V.; SATAROVA, L.M.;  
SU KHUN-GUY [Su Hung-kuei]; KALICHEVA, I.S.; FIRSOVA, L.D.

Further study of the products of iron spallation by  
660 MeV protons. Radiokhimiia 5 no. 6:721-732 '63,  
(MIRA 17:7)

AMAR NAT; LAVROKHINA, A.S.

Separation of lead and thallium from the products of bismuth  
photodisintegration. Radiokhimiia 5 no. 6:752-736 '63.  
(MIRA 17:7)

S/063/63/008/002/015/015  
A057/A126

AUTHORS: Lavrukhina, A.K., Malyshev, V.V., Rodin, S.S.

TITLE: The application of zirconium molybdate and titanium dioxide to the group separation of elements

PERIODICAL: Zhurnal vsesoyuznogo khimicheskogo obshchestva imeni D.I. Mendeleeva, v. 8, no. 2, 1963, 227 - 229

TEXT: The separation of elements by means of ion-exchange columns filled with zirconium molybdate and titanium dioxide was investigated. In the present paper results are given on the separation of the basic fission elements Rb and Cs from Sr and Ba, and from rare earths. Zirconium molybdate was prepared by very slow addition of 200 ml 1.4 M ammonium molybdate solution to 400 ml 1.2 M zirconium chloroxide solution at vigorous stirring, which was continued after the precipitation for 15 min. The precipitate was filtered off, washed for 24 h and dried for 100 h. If suspended in water, 0.2 - 0.5 mm diameter particles were obtained. The same technique was applied to the preparation of titanium dioxide from 200 ml 7% titanium tetrachloride solution and a 20% surplus of 20%

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The application of zirconium molybdate and ....

S/063/63/008/002/015/015  
A057/A126

ammonia solution. The obtained inorganic ion exchange substances were filled into glass columns (5 cm long, 0.5 cm<sup>2</sup> inner cross section), 2 cm high. In preliminary experiments the sorption of Rb<sup>86</sup>, Cs<sup>134</sup>, Fr<sup>212</sup> and Sr<sup>90</sup> was determined by the batch technique using the hydrogen and ammonia form respectively of the exchange substance. Rb, Cs, and Fr did not adsorb on the ammonia form neither from the neutral nor from the 0.3 M NH<sub>4</sub>Cl solution, while Sr adsorbed with 87.5%. From 0.3 M HCl 11.8% Cs, 14.2% Fr, but no Sr was adsorbed by zirconium molybdate. The effect of separation of Sr<sup>90</sup> with 0.1 M HCl from Cs<sup>134</sup> with 4 M NH<sub>4</sub>NO<sub>3</sub> or from Pr<sup>142</sup> with 4 M NH<sub>4</sub>NO<sub>3</sub> on zirconium molybdate in H<sup>+</sup> form is incomplete, since about 10% of the cesium activity remains on the columns. Cs<sup>134</sup> was eluted with 95 - 97% efficiency using as eluent a mixture of 4 M NH<sub>4</sub>NO<sub>3</sub> and 2 M HCl. The method was developed for the separation of short lived radioisotopes and of highly active products respectively. There is 1 figure.

ASSOCIATION: Institut geokhimii i analiticheskoy khimii im. V.I. Vernadskogo  
AN SSSR (Institute of Geochemistry and Analytical Chemistry imeni  
V.I. Vernadskiy AS USSR)

SUBMITTED: May 28, 1962  
Card 2/2



S/048/63/027/001/042/043  
B108/B180

AUTHORS: Lavrukhina, A. K., Moskaleva, L. P., and Kuznetsova, R. I.

TITLE: Some new data on the mechanism of the formation of light nuclei

PERIODICAL: Akademiya nauk SSSR. Izvestiya. Seriya fizicheskaya, v. 27, no. 1, 1963, 137-140

TEXT: Earlier work (Report on the 19. Congress on Pure and Applied Chemistry, Canada, 1961) is continued on the production of light nuclei under bombardment by fast protons. This study covers the energy distribution of  $\text{Na}^{24}$  produced from Al, Ag, and U by bombardment with 660-Mev protons. To establish a dependence on the proton energy the authors also studied the production cross sections and the angular distributions of some light nuclei produced by 120-Mev protons. The production cross sections of  $\text{Be}^7$ ,  $\text{F}^{18}$ ,  $\text{Na}^{24}$ , and  $\text{P}^{32}$  from Sb, Sn, and U have the same course for 660-Mev and 120-Mev protons, but are in the former case higher by about one order of magnitude. The ratio of the

Card 1/2

Some new data on the mechanism ...

S/048/63/027/001/042/043  
B108/B180

yields in  $\text{Na}^{24}$  and  $\text{F}^{18}$  is somewhat lower for 120 than for 660 Mev. For an Sb target it is 1.6, 2.2 for Sn, and 2.1 for U. These data disprove the meson mechanism of momentum transfer in the nucleus. This paper was read at the 12. Annual Conference on Nuclear Spectroscopy, Leningrad, January 26 - February 2, 1962. There are 4 figures and 3 tables.

ASSOCIATION: Institut geokhimii i analiticheskoy khimii im.  
V. I. Vernadskogo Akademii nauk SSSR (Institute of  
Geochemistry and Analytical Chemistry imeni V.I. Vernadskiy  
of the Academy of Sciences USSR)

Card 2/2

MOSKALEVA, L.P.; LAVRUKHINA, A.K.

Angular and energy distributions of  $\text{Na}^{24}$  nuclei emitted during  
the irradiation of aluminum by 660 Mev. protons. Izv. AN SSSR.  
Ser. fiz. 27 no.10:1270-1272 0 '63. (MIRA 16:10)

I 10198-63

EPP(c)/EPP(n)-2/EAT(m)/BDS--APFTC/ASD/3SD--

Pr-4/Pu-4

ACCESSION NR: AP3000029

S/0056/63/044/005/1429/1436

AUTHOR: Lavrukhina, A. K.; Revina, L. D.; Malyshev, V. V.; Satarova, L. M.

TITLE: Spallation of Fe Nuclei induced by 150-MeV protons

SOURCE: Zhurnal eksper. i teoret. fiziki, v. 44, no. 5, 1963, 1429-1436

TOPIC TAGS: Nuclear reactions, iron, low-energy protons, spallation, isotope distribution

ABSTRACT: Continuing their earlier work on the spallation of iron isotopes by 660-MeV protons (Geokhimiya, no. 11, 955, 1961 and Radiokhimiya, in press), the authors studied nuclear reactions at lower energies, aimed at clarifying volume effects in the distribution of cosmogenic nuclides in meteorites. To this end, the main features of spallation of iron nuclei by 150-MeV protons were studied. An empirical equation is found for the production cross sections of the spallation products. The majority of the product nuclei were found to be near the bottom of the stability valley. The weighted numbers of the emitted neutrons and protons are 2.9 and 2.7, respectively. The cross section for the

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L 10198-63  
ACCESSION NR: AP3000029

3

inelastic cross section of 150-MeV protons with iron nuclei is 568 plus or minus 162 mb. The considerable difference between the distributions of the products at 150 and 660 MeV proton energies is probably due to the formation, absorption, and scattering of pions, which increases the probability of transferring large excitation energy to a nucleus at 660 MeV proton energy. Comparison of the total cross section for the inelastic interaction of the iron nuclei with the protons at the two energies with optical-model calculations yields an estimate for the radius of the Fe-56 nucleus, namely  $(1.21) \cdot 10 \sup -13$  cm. The authors express their gratitude to I. S. Kalicheva, L. D. Firsova, and T. I. Kholodkovskaya who took part in this work.

ASSOCIATION: none

SUBMITTED: 06Oct62 DATE ACQ: 12Jun63 ENCL: 00

SUB CODE: PH NR REF SOV: 005 OTHER: 016

bm/CL  
Card 2/2

S/020/63/148/005/010/029  
B102/B186

AUTHORS: Lavrukhina, A. K., Kolesov, G. M.

TITLE: New neutron-deficient isotopes of the cerium group of rare-earth elements

PERIODICAL: Akademiya nauk SSSR. Doklady, v. 148, no. 5, 1963, 1047 - 1048

TEXT: A 99.96%  $\text{Pr}_6\text{O}_{11}$  target was bombarded by 660-Mev protons in the synchrotron of the Ob'yedinennyy institut yadernykh issledovaniy (Joint Institute of Nuclear Research), the rare-earth elements were chromatographically separated and Ce was obtained by extraction, as a radiochemically pure product. The measurements were made with an MTK-20 (MTS-20) end-window counter, and a gamma scintillation spectrometer with a 100-channel pulse-height analyzer. The  $\text{Pr}^{134}$  half-life was determined by separating the 3.1-day  $\text{Ce}^{134}$  from the irradiated praseodymium. From the time dependence of the  $\text{Ce}^{134}$  activity, the  $\text{Pr}^{134}$  half-life was found to be 36 min (mean value  $40 \pm 7$  min). The same period was observed for the 720-kev gamma

Card 1/2

S/020/63/148/005/010/029  
B102/B186

New neutron-deficient ...

component; it possibly arises on  $\text{Pr}^{134}$  decay. From the La decay curve the existence of the following isotopes could be determined:  $\text{La}^{132,133}$  (4.3hrs),  $\text{La}^{131}$  (1.1 hrs) and an isotope with A=129 and a 20-min halflife - probably  $\text{La}^{129}$  with  $\sim 24$  min. The gamma spectrum of the Ce fraction has a  $80 \pm 15$ -kev peak; that of the La fraction peaks at  $115 \pm 20$  kev and  $175 \pm 15$  kev and a 2.2-hr halflife, which could be attributed to  $\text{Ba}^{129}$ . In the Ce fraction also a 13-min activity and gamma peaks at  $80 \pm 15$ ,  $315 \pm 20$  and  $745 \pm 20$  kev ( $\sim 15$  min) were observed. From this the 129-isobar decay series is assumed to be  $\text{Ce}^{129} \xrightarrow{\sim 13\text{min}} \text{La}^{129} \xrightarrow{\sim 20\text{min}} \text{Ba}^{129} \xrightarrow{2.2\text{hr}} \dots \rightarrow \text{Xe}^{129}$ . There are 2 figures.

ASSOCIATION: Institut geokhimii i analiticheskoy khimii im. V. I. Vernadskogo Akademii nauk SSSR (Institute of Geochemistry and Analytical Chemistry imeni V. I. Vernadskiy of the Academy of Sciences USSR)

PRESENTED: August 27, 1962, by A. P. Vinogradov, Academician

SUBMITTED: July 27, 1962

Card 2/2

L 19526-65 EWT(1)/EWT(m)/EWG(v)/FGC/EWA(d)/EEG-4/EEG(t)/EWA(h) Pb-4/  
 Po-4/Fe-5/Pg-4/Pag-2/Peb/Pi-4 DIAAP/ASD(f)-2/AFWL/AFETR/SSD/AFMDC/ESD(t)  
 ACCESSION NR: AP5000/19 GW/WS S/0007/64/000/012/1219/1227

AUTHORS: Lavrukhina, A. K.; Kuznetsova, R. I.; Satarova, L. M.

TITLE: The rate of radioactive isotope<sup>14</sup> formation in chondrites by cosmic rays

SOURCE: Geokhimiya, no. 12, 1964, 1219-1227

TOPIC TAGS: geochemistry, meteorite, cosmic ray

ABSTRACT: The interaction between cosmic rays and nuclei of elements commonly occurring in meteorites has been difficult to study because of the variety of minerals and phases in which these elements are found. Because of this, little information has been gained concerning the rate of isotope production by cosmic radiation in stony meteorites. As an approach to this problem, the authors studied the radioactive products from interaction between protons having energies of 120 and 660 Mev and aluminum nuclei. This element has an atomic weight and an atomic number comparable to the weighted mean values of elements in chondrites (those about which information is desired). Experimental results show that when the numerical difference between atomic numbers of the target element and of the radioactive product is less than 10, the formation cross section remains unchanged in the range of proton energies from 120 Mev to 28 Gev. If the difference is

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L 19626-65

ACCESSION NR: AP5000419

greater than 10, the formation cross section of isotopes increases markedly with increase in proton energies from 120 Mev to 1 Gev. A semi-empirical formula was used to compute these formation cross sections, and the results agree with experimental measurements for a proton energy of 660 Mev. The maximal rate of isotope formation from cosmic rays was computed on the basis of average elemental composition for the center of a chondrite 10 cm in radius. The results are in fair agreement with the observed ratio of decay in the Harleton chondrite, but are only about one-half the values observed for the Bruderheim chondrite. Apparently secondary particles, especially low-energy neutrons, played an important role in the isotope formation of the latter chondrite. These particles were not considered in the computations. Lack of knowledge of exact chemical composition and dimensions of these two chondrites makes it difficult to make any more detailed comparison. Orig. art. has: 4 tables and 6 formulas.

ASSOCIATION: Institut geokhimi i analiticheskoy khimii im. V. I. Vernadskogo, AN SSSR, Moscow (Institute of Geochemistry and Analytical Chemistry, AN SSSR)

SUBMITTED: 13Jul64

ENCL: 00

SUB CODE: AA,ES

NO REF SOV: 009

OTHER: 024

Card 2/2

VONOGRAOV, A.P., akademik; LAVRUKHINA, A.K.; REVINA, L.D.

Nuclear reactions in iron meteorites. Meteoritika no.24:  
22-28 '64. (MIRA 17:5)

LAVRUKHINA, A.K.; KOLESOV, G.M.

• Study of the nuclear disintegration products of rare-earth  
elements. Radiokhimiia 6 no. 1:62-66 '64. (MIRA 17:6)

LAVRUKHINA, A.K.; RODIN, S.S.

Review of Ed. C.E. Crouthamel's book "Progress in nuclear  
energy (Analytical chemistry. Vol.3)." Zhur. anal. khim.  
19 no.3:403 '64. (MIRA 17:9)

L 3858f-65 EWT(m)/EWA(h)

ACCESSION NR: AP5008091

S/0030/65/000/002/0089/0090

AUTHOR: Lavrukina, A. K. (Doctor of chemical sciences)

50  
23

TITLE: Third radiochemical conference

SOURCE: AN SSSR. Vestnik, no. 2, 1965, 89-90

TOPIC TAGS: radiochemistry, radioactive material, radioisotope, radioactive carbon

ABSTRACT: The Third Radiochemical Conference took place in Liblitsy from September 28 to October 1, 1964. It was organized by the Nuclear Chemistry section of the Czechoslovakian Chemical Society with the participation of scientists from Hungary,

Card 1/2

L 38588-65

ACCESSION NR: AP5008091

by the  $C^{14}$  tracer, and the compounds tagged by  $I^{131}$ . X. Beren and G. Nad' reported on the method of  $F^{18}$  separation without a carrier. The second session was dedicated to the use of radioactive isotopes in technology. O. Shtepan discussed the currents during glass melting.

"APPROVED FOR RELEASE: 06/20/2000

CIA-RDP86-00513R000928830009-2

microquantities in the presence of heavy

ASSOCIATION: none

SUB CODE: NP

SUBMITTED: 00

ENCL: 00

NO REF SOV: 000

OTHER: 000

Card 2/2

APPROVED FOR RELEASE: 06/20/2000

CIA-RDP86-00513R000928830009-2"

LAVRUKHINA, A.K.; KUZNETSOVA, R.I.; SATAROVA, L.M.

Formation rate of radioactive isotopes in chondrites under the  
action of cosmic rays. Geokhimiia no.12:1219-1227 D 1964. (MIRA 18:8)

1. Institut geokhimii i analiticheskoy khimii imeni V.I.Vernadskogo  
AN SSSR, Moskva.



L 7696-66 EWT(1)/EWA(d) GW  
ACC NR: AP5028896

SOURCE CODE: UR/0011/65/000/012/0003/0011

AUTHOR: Lavrukhina, A. K. 44,55

ORG: Institute of Geochemistry and Analytical Chemistry im. V. I. Vernadskiy, 44,55  
AN SSSR (Institut geokhimii i analiticheskoy khimii AN SSSR) B

TITLE: Investigating the evolution of meteoritic matter from the effects of the  
nuclear reactions induced by cosmic rays 44,55 12

SOURCE: AN SSSR. Izvestiya. Seriya geologicheskaya, no. 12, 1965, 3-11

TOPIC TAGS: meteorite, asteroidal body, chondrite, geochemistry, cosmogony,  
meteoritic matter, cosmic ray, radioactive dating, meteoritics

ABSTRACT: Current theories on the origin, age, and composition of meteorites are evaluated against a background of new data obtained from orbital studies, geochemical analysis, and radioactive methods of age determination. Thus, for example, V. G. Fesenkov has determined that the orbit of the Sikhote-Alin' meteorite was typically asteroidal, while E. Anders, on the basis of orbital studies, has traced the parent bodies of meteorites to 34 asteroids intersecting the orbit of Mars. A. A. Yavnel' has developed a meteorite classification system, based on the nickel content of the meteoritic metal, that indicates a close genetic relationship between all meteorites and suggests a common origin. Three model theories of meteorite formation and evolution are discussed: 1) the Urey model in which achondrites and iron meteorites are

Cord 1/3

UDC: 522.6

0701.2114

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ACC NR: AP5028896

believed to have been formed in primary bodies of lunar dimensions, while chondrites were formed in secondary bodies of asteroidal size from the ducts of disintegration of the primary bodies; 2) the model, attributed to Anders and Ringwood, involving the differentiation of matter from a parent body; and 3) the model, attributed to Vinogradov and Wood, in which chondrules are believed to have developed from a protoplanetary dust cloud. Examination of age data obtained by methods based on the decay of all natural radioactive elements ( $U^{238}$ ,  $Re^{187}$ ,  $K^{40}$ ,  $Rb^{87}$ ) indicates that meteoritic matter solidified some 4.5 million years ago. In the light of the hypothesized common process of formation of all bodies of the solar system from the protoplanetary cloud, this determination of the time of solidification of the matter of the solar system is considered to be a major advance in meteoritics. Data now available indicate three possible processes whereby meteorites developed into separate astronomical bodies: 1) Larger bodies disintegrated into progressively smaller meteoritic bodies. 2) Meteorites falling onto the earth were actually formed on the moon when other meteorites originating in the asteroid belt collided with the earth's satellite. In this process the surface layer of the moon provided the matter found in so-called hypersthene meteorites, while matter from the lunar interior entered into the composition of so-called bronzite meteorites. 3) Hypersthene meteorites originated in the collision and subsequent disintegration of large asteroids some 500,000,000 years ago. It is concluded, however, that all available data on bronzite meteorites attest to their formation about 5,000,000 years ago in the disintegration of a single primary body whose matter did not undergo the changes evident in hypersthene meteorites. This is in accord with the determination of a higher content of metallic iron and lower

Cord 2/3

L 7696-66

ACC NR: AP5028896

content of oxidized iron in bronzites than in hypersthénics. Orig. art. has: 6 fig-  
ures. [DM]

SUB CODE: AA/ SUBM DATE: 29Mar65/ ORIG REF: 013/ OTH REF: 012/ ATD PRESS: 4141

Card 3/3

LAVRUKHINA, A.K.

Meteorites and their use in studying cosmic rays. Izv. AN SSSR.  
Ser.fiz. 29 no.10:1838-1842 0 '65. (MIRA 18:10)

1. Institut geokhimii i analiticheskoy khimii im. V.I.Vernadskogo  
AN SSSR.

LAVRUKHINA, Augusta Konstantinovna; KOLESOV, Gennadiy Mikhaylovich;  
KALYUZHNAYA, T.P., red.

[Isotopes in the universe] Izotopy vo Vselennoi. Moskva,  
Atomizdat, 1965. 239 p. (MIRA 18:8)

LAVRUKHINA, A.K.; RUTKOVSKIY, V.M.; IERAYEV, T.A.; YUKINA, I.V.

Study of the variations in cosmic rays based on their effects on  
stony meteorites. Izv. AN SSSR. Ser. fiz. 29 no. 10:1849-1855 1985.  
(MIRA 18:10)

1. Institut geokhimii i anorganicheskoy khimii im. V.I. Vernadskogo  
AN SSSR.

GOL'DANSKIY, V.I., otv. red.; LAVRUKHINA, A.K., prof., doktor  
khim. nauk, otv. red.; RODIN, S.S., red.; PROKOP'YEV, Ye.P.,  
red.

[Nuclear chemistry] IAdernaia khimiia. Moskva, Nauka, 1965.  
327 p. (MIRA 18:12)

1. Akademiya nauk SSSR. Institut geokhimi i analiticheskoy  
khimi. 2. Chlen-korrespondent AN SSSR (for Gol'danskiy).

LAVRUKHINA, A.K., prof.

Meteorites and space chemistry. Priroda 55 no.1:33-44  
Ja '66. (MIRA 19:1)

1. Institut geokhimii i analiticheskoy khimii im. V.I.  
Vernadskogo AN SSSR, Moskva.



L 39933-66 EWT(1)/EWT(m) GW/GD

ACC NR: AT6017648

SOURCE CODE: UR/0000/65/000/000/0007/0073

AUTHOR: Lavrukhina, A. K.

ORG: none

19  
TITLE: Effect of nuclear reactions caused by fast protons in meteorites 53  
B+1

SOURCE: AN SSSR. Institut geokhimii i analiticheskoy khimii. Yadernaya khimiya  
(Nuclear chemistry). Moscow, Izd-vo Nauka, 1965, 7-73

TOPIC TAGS: nuclear reaction, fast particle, meteorite, radioactive decay, cosmic ray, geochemistry, geophysics, cosmology

ABSTRACT: An extensive survey of the new field of nuclear space chemistry, concerned mainly with the stable and radioactive isotopes found in iron and stony meteorites is presented. The author discusses the basic problems of discovering the principles governing the origin and distribution of cosmogenic isotopes in planets, satellites, asteroids, meteorites, and cosmic dust and of explaining their role in the formation and evolution of these bodies. Such studies are heavily based on the recent development of high-sensitivity radiometers of extremely low background and mass-spectrometers for the study of microquantities of stable and radioactive cosmogenic isotopes. The survey embraces the following topics: chemical composition of meteorites, isotopic composition of the elements in meteorites, variations in the content of certain

Card 1/2

L 39933-66

ACC NR: AT6017648

stable cosmogenic isotopes in meteorites, cosmogenic radioisotopes, nuclear reactions in meteorites, methods of determining the rates of formation of cosmogenic isotopes in meteorites, rates of formation of cosmogenic isotopes in iron meteorites, rates of formation of cosmogenic isotopes in stony meteorites. Orig. art. has: 26 tables, 13 figures.

SUB CODE: 03/20/88/ SUBM DATE: 03Nov65/ ORIG REF: 052/ OTH REF: 291

Card 2/2

L 40244-66 EWT(1)/EWT(m)/FCC/T/EWP(t)/ETI IJP(c) GW/JD

ACC NR: AT6020807

SOURCE CODE: UR/2534/65/000/026/0091/0101

AUTHOR: Lavrukhina, A. K.

ORG: none

TITLE: Study of space and time variation of cosmic rays based on the effects of nuclear fission in meteorites

SOURCE: AN SSSR. Komitet po meteoritam. Meteoritika, no. 26, 1965, 91-101

TOPIC TAGS: cosmic ray, cosmic ray intensity, meteorite, nuclear fission, SOLAR ACTIVITY

ABSTRACT: In this article the author discusses the spatial distribution of cosmic rays in the solar system, the time variations of the intensity of cosmic rays, and the intensity of cosmic rays along meteoritic orbits. The information obtained by means of meteorites concerning the intensity of cosmic rays permits the conclusion that the intensity of cosmic rays changes depending upon the intensity of solar activity even at appreciable distances from the sun, in any case at a distance of several astronomical units and that the intensity of cosmic rays is approximately by a factor of 2 higher in the region of meteoritic orbits in comparison with that in the vicinity of the earth. The data presented on the intensity of cosmic rays and their spatial distribution in the solar system are undoubtedly still insufficiently accurate and incom-

Card 1/2

SHEYNIN, Yulian Mikhaylovich; RUBINSHTEYN, M.I., doktor ekon. nauk,  
otv. red.; LAVRUKHINA, I.M., red.; ASTAF'YEVA, G.A.,  
tekhn. red.

[Science and militarism in the U.S.A.; scientific and  
technological revolution in military art and the origina-  
tion of conditions for the crisis of militarism] Nauka i  
militarizm v SShA; nauchno-tekhnicheskii perevorot v voen-  
nom dele i vozniknovenie predposylok krizisa militarizma.  
Moskva, Izd-vo AN SSSR, 1963. 590 p. (MIRA 16:12)  
(United States--Military art and science)  
(United States--Militarism)

ANTONOV, E.I., inzh.; KUZNETSOV, D.P., inzh.; LAVRUKHINA, T.P., inzh.;  
TSYRKIN, I.Z., inzh.

Redesigning of the EP-3-600 ejector for operation on steam pressures  
of 6 atm. Energetik 10 no.5:13-16 My '62. (MIRA 15:5)  
(Steam turbines)

LA VRUSEVICH, A.I.

New genus *Chavsakia* from Indlow sediments in the Zeravshan-Gissar mountain region. Izv. Otd. est. nauk AN Tadzh. SSR no.1:35-41 '59.  
(MIRA 13:3)

1. Institut geologii AN Tadzhikskoy SSR.  
(Arg Valley--Corals, Fossil)

LAVRUSEVICH, A.I.

Holomophyllum with a scaly epitheca from Silurian sediments of the  
Zeravshan and Gissar Ranges. Dokl. AN Tadzh. SSR 3 no. 4: 21-25. '60.  
(MIRA 14:4)

1. Institut geologii AN Tadzhikskoy SSR. Predstavleno akademikom  
AN Tadzhikskoy SSR A.P. Nedzvetskim.  
(Tien Shan—Corals, Fossil)

LELESHUS, V.L.; MENAKOVA, G.N.; LAVRUSEVICH, A.I.

Silurian stratigraphy of the southern Tien Shan. Dokl.AN  
SSSR 133 no.1:196-198 J1 '60. (MIRA 13:7)  
(Daurich region--Geology, Stratigraphic)



LAVRUSEVICH, A.I.; MUCHAIDZE, D.R.

Stratigraphy of the calcareous shale formation of the eastern part  
of the Zeravshan-Gissar Range. Izv. Otd. geol.-khim. i tekhn. nauk  
AN Tadzh.SSR 1:61-67 '60. (MIRA 15:1)

1. Upravleniye geologii i okhrany nedr pri Sovete Ministrov  
Tadzhikskoy SSR.

(Zeravshan Valley--Shale)

LAVRUSEVICH, V., laureat Stalinskoy premii.

Universal metal shoring for dock work. Mor. i rech.flot 14 no.5:  
27-29 My '54. (MLRA 7:7)  
(Dry-docks)

ZAGLUBOTSKIY, P.M.; DURNOV, G.P.; LAVHUSEVICH, V.V.; MIKHAYLENKO, V.I.;  
IVANOV, V.M., spetsred.; SHUIN, V.I., red.; FORMALINA, Ye.A.,  
tekh.n.red.

[Practices of efficiency promoters in ship repairing] Opyt  
ratsionalizatorov v sudoremonte. Moskva, 1959. 53 p.  
(MIRA 13:9)

(Ships---Maintenance and repair)

UGOLEVA, N.A., BESKINA, S.R., LAVRUSHENKO, V.A.,

"Biochemical and histochemical studies of nuclei acids in chick  
embryo chorioallantoic membrane infected with Sendai virus.

Report submitted to the Intl. Congress for Microbiology  
Montreal, Canada 19-25 Aug 1962

LAVRUSHENKOVA, Z.A., Cand Med Sci -- (diss) "<sup>Accessory</sup>~~Additional~~

sinuses of the nose .in the normal and pathological

case <sup>under</sup>~~in~~ X-ray <sup>exposure</sup>~~lighting~~." Smolensk, 1958, 17 pp

(Second Mos State Med Inst im N.I. Pirogov) 250 copies

(KL, 32-58, 112)

- 70 -

LAVRUSHENKOVA, Z. A.

USSR / Human and Animal Morphology (Normal and Pathological).  
Skeleton.

S

Abs Jour : Ref Zhur - Biol., No 21, 1958, No 97137

Author : Lavrushenkova, Z.A.

Inst : Smolensk Medical Institute

Title : Variations of Frontal Pauses According to Data of  
Roentgenological Investigation.

Orig Pub : Tr. Smolenskogo med. in-ta, 1957, 6, 291-296

Abstract : On 1,154 roentgonograms of paranasal pauses of sinuses of humans aged from 16-77 years (510 males and 644 females), the following classification of frontal sinuses (FS) was established: absence of both FS (2.34% of cases), unilateral development of FS (3.72%), solitary (1.37%), double cavitation (69.5%), triple cavitation (18.29%), quadruple cavitation (3.81%), quintuple cavitation (0.44%), multiple cavitation of FS (0.17%). Within the limits of these groups, more detailed subdivisions are singled out.

Card 1/1

ALEKHIN; BORISOV; VOLKOV; GRIGOR'YANTS; GRUZDEV; DICH; DUSEYEVA;  
LAVRUSHIN; LOPINSKIY; IVANOVA;; KONKIN; MEOS; MIKHAYLOV;  
MOGILEVSKIY; PAKSHVER; ROGOVIN; TAIROV; SHIFRIN

Deserving workers of the synthetic fibers industry. Khim.  
volok. no.3:79 '61. (MIRA 14:6)  
(Birger, Georgii Efimovich, 1886)

L 40244-66

ACC NR: AT6020807

plete, and it is necessary to conduct further detailed investigations to explain the character of the distribution of cosmogenic isotopes in relation to the chemical composition and size of meteorites and to the depth of occurrence of the investigated specimens in the meteorite. Orig. art. has: 5 tables, 7 figures, and 1 formula.

SUB CODE: 03,04,18/ SUBM DATE: 00/ ORIG REF: 001/ OTH REF: 032

Card 2/2 MLP



L 37000-66 EWT(1) GW

ACC NR: AP6023200

SOURCE CODE: UR/0020/66/168/006/1275/1278

50  
47  
B

AUTHOR: Lavrukhina, A. K.

ORG: Institute of Geochemistry and Analytical Chemistry im. Vernadskiy,  
AN SSSR (Institut geokhimii i analiticheskoy khimii AN SSSR)

TITLE: Determination of preatmospheric meteoritic dimensions

SOURCE: AN SSSR. Doklady, v. 168, no. 6, 1966, 1275-1278

TOPIC TAGS: meteoric size, terrestrial atmosphere, cosmogenetic isotope, cosmic ray, excitation function, nuclear reaction, proton energy, secondary nuclear particle, *METEOR, COSMIC RAY EFFECT, COSMIC DUST*

ABSTRACT: Meteors are the source of cosmic matter. Meteoric size before entry into the terrestrial atmosphere is unknown. The only way to determine the size of meteors before they enter the terrestrial atmosphere is to study the formation speed of cosmogenetic isotopes generated by the interaction of cosmic rays with the meteoric matter. This method is based on the difference in excitation functions of various nuclear-reaction product groups. Nuclear-reaction products are classified into two groups with  $\Delta A < 10$  and  $\Delta A > 10$ , where  $\Delta A$  is the difference of products gained by protons of different energies. The speed of isotope  $H_{\Delta 1}$  formation in the i-point in the meteor body is a

Card 1/2

UDC: 523.51+539.17+523.165

L 37000-66

ACC NR: AP6023200

3  
function of the intensities of primary cosmic radiation and active secondary nuclear particles. The particle spectrum and the energy function determine the role of each kind of particle in the  $A_i$  isotope formation. Isotope formation in meteors of various radii is distributed differently, depending upon the length of the radius and the production class group. The quantity of secondary mesons and protons was determined from their spectra as obtained in the atmosphere. Nomograms were drawn for isotopes from the data of a series of meteoric bodies. The accuracy of the author's method is based on radiometric determination of low cosmic-ray activity and their stability in the solar system. The author expresses thanks to L. D. Revina, T. A. Ibrayev and T. I. Kholodkovskaya for help. Orig. art. has: 1 table, 3 figures, and 2 formulas. [EG]

SUB CODE: 03/ SUBM DATE: 25Aug65/ ORIG REF: 005/ OTH REF: 009  
ATD PRESS: 5035

Card 2/2 *JS*

L 40347-66 EMT(m)/EWP(t)/ETI IJP(c) JD

ACC NR: AP6019436

(A)

SOURCE CODE: UR/0007/66/000/003/0281/0290

AUTHOR: Lavrukhina, A. K.; Kolesov, G. M.; Kalicheva, I. S.; Akol'zina, L. D.

ORG: Institute of Geochemistry and Analytical Chemistry im. V. I. Vernadskiy, AN SSSR,  
Moscow (Institut geokhimii i analiticheskoy khimii AN SSSR)

TITLE: Activation determination of Ce, Eu, Sc, Ba, U, and P in dark and clear varieties of Kunashak and Pervomayskiy village chondrites

SOURCE: Geokhimiya, no. 3, 1966, 281-290

TOPIC TAGS: neutron activation analysis, meteorite, radioisotope, cerium, europium, scandium, barium, uranium, phosphorus

ABSTRACT: <sup>19</sup> Neutron activation analysis was used to determine various elements in Kunashak and Pervomayskiy village chondrites. The samples were irradiated with a neutron flux of  $1.2 \times 10^{13}$  n/cm<sup>2</sup> sec. The content of Eu, Sc, Ba, and U in the clear varieties of chondrites of olivine-hypersthene composition generally correspond to the average content previously found in chondrites of this type. The concentration of Eu, Ce, and Sc in hypersthene chondrites is higher than in enstatite chondrites. In nonmagnetic fractions of the investigated meteorites as compared to their unseparated samples, the content of Ce, Eu, and Sc is higher, owing to the lithophilous character of these elements. The concentration of Eu, Sc, Ba, and U in the dark and clear varieties of Kunashak meteorite is approximately the same. The P concentration in the dark varieties of

Card 1/2

UDC: 550.42+552.6

L 40347-66

ACC NR: AP6019436

7

chondrites is higher than in the clear ones. Data on cerium are of particular interest: the Ce content in clear varieties is about twice that in dark ones, which correlates with the lower concentration of metallic iron and higher concentration of ferrous iron, manganese, and chromium in clear varieties for the same total content of iron and troilite in both varieties. These data and also data on the content of inert gases indicate that the substance of the clear variety of the chondrites studied had undergone a more extensive oxidation than the substance of the dark variety, i. e., that the two varieties had a different thermal history. Authors thank T. F. Yakubova for assistance in the measurement of the radioisotopes, Yu. V. Yakovlev, N. N. Dogadkin, and A. Z. Miklishanskiy for placing the samples in the reactor, and V. Ya. Kharitonova and M. I. D'yakonova, on the staff of the Committee on Meteorites, AN SSSR (Komitet po meteoritam AN SSSR) for providing the meteorite samples. Orig. art. has: 6 figures and 4 tables.

SUB CODE: 03,07/ SUBM DATE: 31Jul65/ ORIG REF: 016/ OTH REF: 014

*ms*  
Card 2/2

CHUYANOV, A., inzhener; OL'SHANSKIY, I., inzhener; LAVRUSHIN, A., inzhener.

The leader in the meat packing industry; twentieth anniversary of the  
Moscow Meat Combine. Mias. ind. SSSR 24 no. 6:5-8 '53. (MIRA 6:12)  
(Moscow--Meat industry) (Meat industry--Moscow)

LAVRUSHIN, A., inzhener; OL'SHANSKIY, I., inzhener.

Re-equipping shops of the Moscow meat combine. Mias.ind. SSSR.  
25 no.5:12-17 '54. (MIRA 7:11)  
(Packing houses)

LAVRUSHIN, A., inzhener; OL'SHANSKIY, I., inzhener

New efficiency methods in the Moscow meat combine. Mias.  
ind. SSSR 26 no.3:53-55 '55. (MLRA 8:9)  
(Moscow--Meat industry)

**LAVRUSHIN, A.**

**We are installing new lines and automatic machinery. Mias.ind.  
SSSR 26 no.4:24 '55. (MIRA 8:10)**

- 1. Moskovskiy myasokombinat imeni A.I.Mikoyana  
(Moscow--Packing houses--Equipment and supplies)**



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